# WIMP Dark Matter Searches with Liquid Xenon and Liquid Argon Detectors

Elena Aprile Columbia University TeVPA08 Beijing, September 27, 2008

# **Principle of Direct Detection**



- Elastic collisions with nuclei
- The recoil energy is:

$$\boldsymbol{E}_{\boldsymbol{R}} = \frac{\left| \boldsymbol{\vec{q}} \right|^2}{2\boldsymbol{m}_N} = \frac{\mu^2 v^2}{\boldsymbol{m}_N} (1 - \cos\theta) \le 50 \ \boldsymbol{keV}$$

• and the expected rate:

$$\boldsymbol{R} \propto \boldsymbol{N} \frac{\boldsymbol{\rho}_{\chi}}{\boldsymbol{m}_{\chi}} \left\langle \boldsymbol{\sigma}_{\chi N} \right\rangle \qquad \mu = \frac{\boldsymbol{m}_{\chi} \boldsymbol{m}_{N}}{\boldsymbol{m}_{\chi} + \boldsymbol{m}_{N}}$$

$$\begin{split} N &= number \ of \ target \ nuclei \ in \ detector \\ \rho_{\chi} &= local \ WIMP \ density, \ m_{\chi} &= WIMP \ mass \\ <\sigma_{\chi N} > &= \ scattering \ cross \ section \end{split}$$

# **Predicted Rates in Different Targets**



#### **Challenges for Dark Matter Direct Detection**

- →large mass (ton scale)
- →low energy threshold (a few keV)
- background suppression
  - deep underground
  - passive/active shielding
  - Iow intrinsic radioactivity
  - >gamma background discrimination

# **Direct Detection Experiments**



# World Wide Experiments for Dark Matter Direct Detection





En la



Soudan

CDMS

LUX



Frejus/ Modane EDELWEISS



Gran Sasso CRESST DAMA/LIBRA WARP XENON

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Boulby

ZEPLIN

DRIFT

# **The Noble Liquids Revolution**

Noble liquids are relatively inexpensive, easy to obtain, and dense

Easily purified

- most impurities freeze out
- low surface binding
- purification easiest for colder liquids

Ionization electrons and scintillation photons provide dual and complementary information on particle energy and ID plus 3D localization when used in TPC

- High ionization (W  $_{LXe}$ = 15.6 eV and W  $_{LAr}$ = 23.6 eV)
- Very high Scintillation yield (~ 40,000 photons/MeV)
- Transparent to their own scintillation
- modest quenching factors for nuclear recoils
- high electron mobility, low e-diffusion

Well suited for large, homogeneous targets and detectors

#### Liquified Noble Gases: Basic Properties

Dense and homogeneous Do not attach electrons, heavier noble gases give high electron mobility Easy to purify (especially lighter noble gases) Inert, not flammable, very good dielectrics Bright scintillators

	Liquid density (g/cc)	Boiling point at 1 bar (K)	Electron mobility (cm <sup>2</sup> /Vs)	Scintillation wavelength (nm)	Scintillation yield (photons/MeV)	Long-lived radioactive isotopes	Triplet molecule lifetime (µs)
LHe	0.145	4.2	low	80	19,000	none	13,000,000
LNe	1.2	27.1	low	78	30,000	none	15
LAr	1.4	87.3	400	125	40,000	<sup>39</sup> Ar, <sup>42</sup> Ar	1.6
LKr	2.4	120	1200	150	25,000	<sup>81</sup> Kr, <sup>85</sup> Kr	0.09
LXe	3.0	165	2200	175	42,000	<sup>136</sup> Xe	0.03

# LXe and LAr for Dark Matter Detection

**complementary nuclei** for SI and SD interactions

scalability : large volume detectors at reasonable cost and technical challenge

Multiple and effective Background reduction schemes: liquid self-shielding, event localization in 3D TPC, event topology

intrinsic background: radioactive Kr-85 in Xe can be reduced to ppt level; radioactive Ar-39 in Ar can be reduced by isotope separation and/or extraction from UG wells

recoil energy threshold : high scintillation yields and continuing progress in photosensors with high QE and low back for both LXe and LAr

gamma rejection: via charge/light ratio and via pulse shape discrimination of light pulse(easier for LAr light)



# Two Phase Xe/Ar Time Projection Chamber (XENON,ZEPLIN,WArP,ArDM)



# **1st Generation Noble Liquid Experiments**

- ZEPLINI, II, III, XENON10, WArP 2I: fiducial mass in the range of ~10 kg.
- Focus review on next generation (2008-2013): fiducial mass in the range 50 1000 kg.

XENON10 (Gran Sasso Lab): 22 kg LXe 2-phase TPC

ZEPLIN II (Boulby Mine): 30 kg LXe 2-phase TPC



WARP (Gran Sasso): New 2.31 LAr 2-phase TPC



#### **XENON10: Best WIMP-Nucleon Cross Section Upper Limits**



## The XENON Dark Matter Phased Program XENON1t



# XENON100 @ LNGS

1.4 km of rock (3100 mwe)muon intensity is approximately 1/(m<sup>2</sup> h)







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12 Ph.D. Students and 10 Postdocs

#### **XENONIOO:** The TPC Assembly

- x10 fiducial mass of XENON10
- x100 less back than XENON10





#### **XENON100: The PMTs**

- 242 PMTs (Hamamatsu R8520-06-AI)
- 1 " square metal channel developed for XENON
- Low radioactivity (<1 mBq U/Th per PMT)
- 80 PMTs for bottom array (33% QE)
- 98 PMTs for top array (23% QE)
- 64 PMTs for top/bottom/side Veto (23% QE)



# 

#### PMTs for Side & Bottom Shield

#### **PMT Base**

#### **XENON100:** Keeping it Clean!



#### Reduce <sup>85</sup>Kr < 50 ppt

Kr85 (Beta, Emax = 687 keV, t = 10.8 y, br = 99.563%) -> Rb85

Kr85 (Beta, Emax = 173 keV, t = 10.8 y, br = 0.434%) -> Rb85m (Gamma, E = 514 keV, t = 2.43 us) -> Rb85

A Cryogenic Distillation Tower is used to reduce by  $10^3$  the initial Kr concentration

XENON100 data used to measure initial concentration (7+-2 ppb)

XENON100 science goal: ~50 ppt

#### Reduce $O_2 < I ppb$

Electronegative impurities reduce charge and light yield.

XENON100 drift is over 30 cm of LXe

Continuous purification by hot getter is used to achieve goal of < 1 ppb O<sub>2</sub>



#### **XENONI00: Projected Sensitivity**



Sensitivity reach is  $\sigma_{SI} \sim 2 \times 10^{-45} \text{ cm}^2$  after 7 months data with zero BKG

Two years of data would give ~20 WIMP events if  $\sigma_{SI}$ ~2x10<sup>-44</sup> cm<sup>2</sup> for 100 GeV

# **Time Evolution of the XENON Program**



# **ZEPLIN III @ Boulby**

- I4 kg (6.6 kg fiducial) dual-phase XeTPC with 31 PMTs in the liquid
- > 3.5 cm drift with high field (4 kV/cm) for better ER/NR
- open plan-no surfaces; all Cu construction



ER gamma band from FSR\_10pc



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# **ZEPLIN III Future Plan**

- Extended gamma calibration run
- Repeat neutron calibration
- Replace PMTs with low activity type
- Predict 30 x background reduction





# LUX (100 kg)@ Homestake

- > 300 kg (100 kg fiducial) dual-phase XeTPC with 122 PMTs
- > 50 kg prototype with 4 PMTs (R8778) being assembled and tested at CWRU
- Plan to move and test full detector above ground at Homestake by end 2008
- > water shield to be built in Davis cavern (once water is pumped from that mine)
- > aggressive background and sensitivity projections (7 x 10<sup>-46</sup> cm<sup>2</sup> after 1 yr)





# XMASS @ Kamioka

- Single phase liq. Xe
  - Pentakis-dodecahedron ← 12 pentagonal pyramids Each pyramid  $\leftarrow$  5 triangles
  - Radius: 39~42 cm
  - 642 Hex. PMT immersed in liq. Xe
  - PMT photo-cathode coverage: 64%
  - Inner mass of Xenon: 857kg (~3g/cm<sup>3</sup>)





#### **New Experimental Hall**



# XMASS (100 kg) @ Kamioka

## Aim:

- 10<sup>-4</sup> dru (ev/kg/keV/day) before any PSD applied
- ➔ 10<sup>-45</sup>cm<sup>2</sup> SI for ~100GeV WIMPs

## For the signal detection

- Low threshold: ~ 5 keV or less
- ← 8 pe /keV (64% photo-cov.)
- Large fiducial volume: 100kg or more

20cmφ: 100kg (20cm self-shield) 25cmφ: 200kg (15cm self-shield)

XMASS (100kg, 5yr) 10<sup>-45</sup>cm<sup>2</sup> (10<sup>-9</sup>pb) for SI and 10<sup>-39</sup>cm<sup>2</sup> (10<sup>-3</sup>pb) for SD



# WArP (100 liter) @ LNGS

- > 140 kg dual-phase ArTPC under construction at LNGS with 10<sup>-45</sup> cm<sup>2</sup> sensitivity reach
- electron recoil discrimination by S2/S1 and by PSD (5ns and 1.6 micros decay times)
- target surrounded by an active LAr shield (9 tons and 300 PMTs)
- electron recoil discrimination by S2/S1 and by PSD of scintillation light, plus event localization
- > Ar gas with low level of radioactive <sup>39</sup>Ar from US u.g. reservoir promising for multi-ton expt



# WArP (2 liter Prototype) @ LNGS

- proof of principle demonstration and 1st DM search with LAr
- exposure of 96.5 kg x days (Jan April 2006): placed a limit at 10<sup>-42</sup>
- studied discrimination power from PSD at 10<sup>-8</sup>
- > new version of 2.3 liter detector installed underground
- > new PMTs, gas recirculation, low activity SS (~ 2 x less back)
- factor of 20 reduction with depleted <sup>39</sup>Ar depleted argon





# ArDM @ LSM

- > 850 kg dual-phase ArTPC under construction at CERN for deployment at Canfranc Lab
- electron recoil discrimination by S2/S1 and by PSD of scintillation light, plus event localization
- expected sensitivity: I WIMP event/ton/day at 10<sup>-44</sup> cm<sup>2</sup>



# **Pulse Shape Discrimination in LAr**



Axion/Wimp WS, DESY 19.6.08 C. Regenfus

ArDM

# **ArDM Status**

- Cryostat and liquid purification and recirculation built and tested
- HV generator (Greinacher circuit) to reach ≈ 4 kV/cm (Vtot = 500 kV) placed in liquid
- Slow Control has been implemented
- Double phase LEMs successfully operated in double phase Ar mode (stable gain of 10<sup>4</sup>); with final LEM charge readout segmented; A/D conversion and DAQ system being developed
- 14 bialkali 8" PMTs (TPB coated) installed at the bottom
- 15 light reflector/shifter foils produced and installed

Amsier et al., "Luminescence quenching of the triplet cimer state by air traces in gaseous argon<sup>a</sup> arXiv:0708.2521



Double-sided copper-clad (35 µm layer) G-10 plates

- Precision holes by drilling
- Palladium deposition on Cu (<~ 1 µm layer) to avoid oxid za</li>
- Single LEM Thickness: 1.5 mm
- Amplification hole diameter = 500 µm
- Distance between centers of neighboring holes = 800 µr





# SUMMARY

- Direct detection experiments with noble liquids have advanced rapidly in past few years. They will play a vital complementary role to indirect and collider searches.
- Sensitivity at the 10<sup>-45</sup> cm<sup>2</sup> (SI) and 10<sup>-39</sup> cm<sup>2</sup> (SD) level within the next couple of years. Great discovery potential!
- Systematic effort to lower the backgrounds while simultaneously increasing target mass is paying off quickly. We are currently seeing gamma background rates down from 1 to 10<sup>-4</sup> cts/keV/kg/day, competitive with the best low background searches!
- The power of a 3D TPC for background identification and suppression and the self shielding property of LXe well established by XENON, ZEPLIN and XMASS prototypes. The larger sensitive volume detectors being deployed or under development today will show a clear advantage for the identification of neutrons by their multiple interactions
- Multiple targets and even same target competition vital for the field at this stage. Natural selection however will occur as scale, cost and complexity increase